

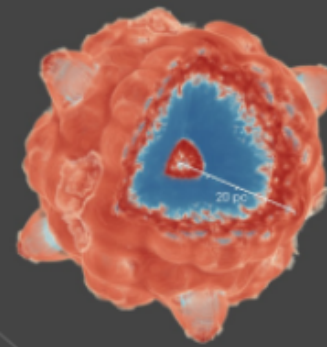
Cloud Formation by Supernova Implosion

Leonard Romano, Manuel Behrendt, Andreas Burkert

New characterization of

radiative stage &

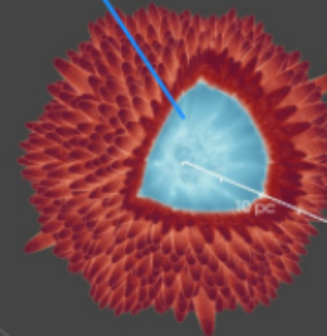
merging with ISM



d) Cloud formation
($t \gtrsim \text{few Myr}$)

$$M_{\text{cloud}} \sim 10^3 - 10^4 M_{\odot}$$

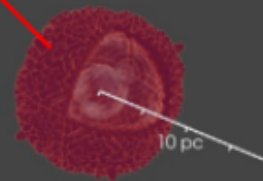
Inflowing Gas



c) Implosion
($t \lesssim \text{few Myr}$)

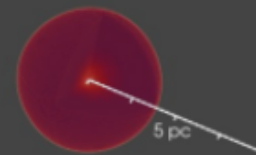
$$P_{\text{shell}} \sim P_{\text{ISM}}$$
$$\dot{M}_{\text{in}} > 0$$

Outflowing Gas



b) Momentum Conserving Snowplow
($t \lesssim 20 - 100 t_{\text{sf}}$)

$$P_{\text{shell}} \gg P_{\text{ISM}}, P_{\text{Bubble}}$$
$$M_{\text{Bubble}} \sim 0$$



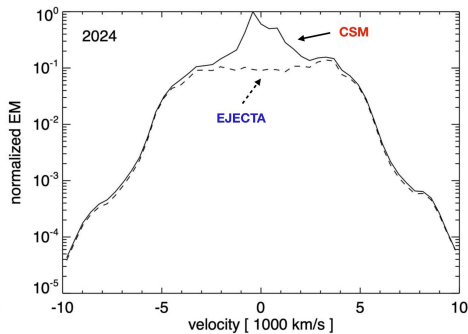
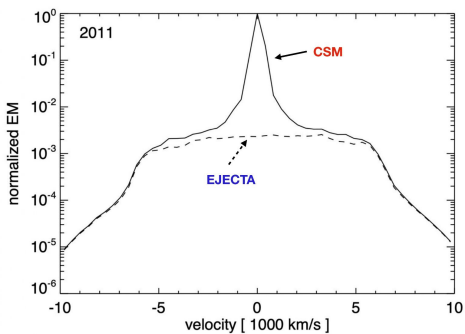
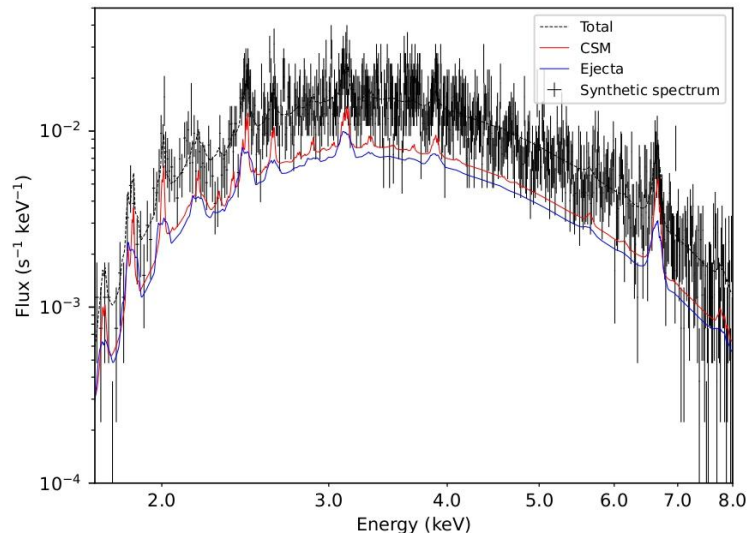
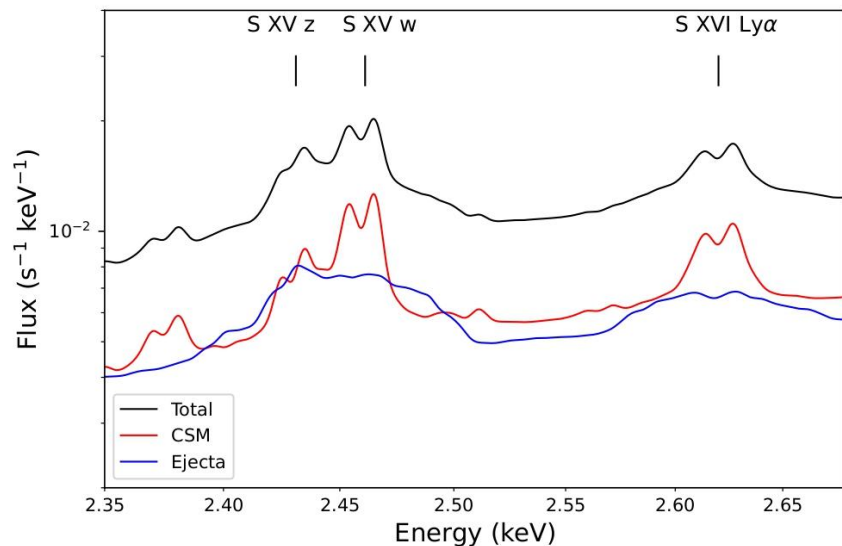
a) Pressure Driven Snowplow
($t < 3 - 5 t_{\text{sf}}$)

$$P_{\text{Bubble}} \gg P_{\text{shell}} \gg P_{\text{ISM}}$$
$$\dot{M}_{\text{Bubble}} < 0$$

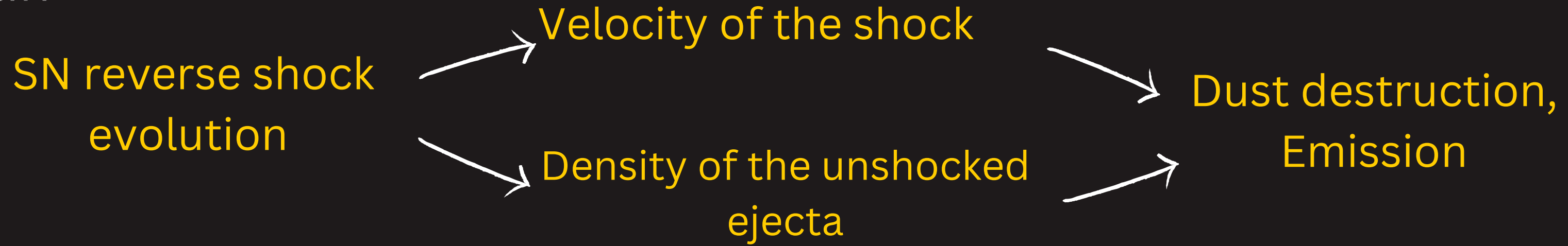


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Probing Shocked Ejecta in SN 1987A: A novel diagnostic approach using XRISM-Resolve (S4.22)

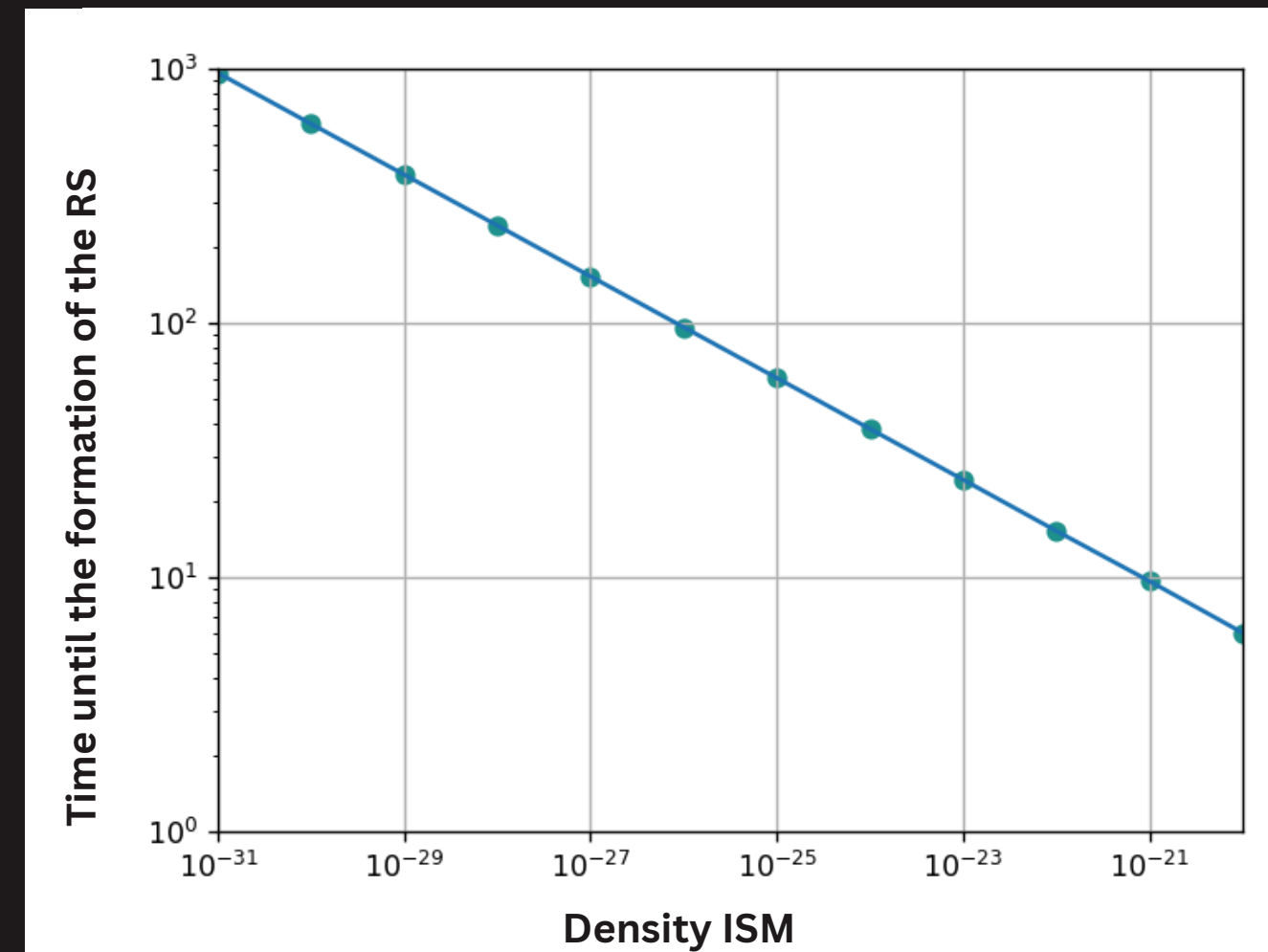
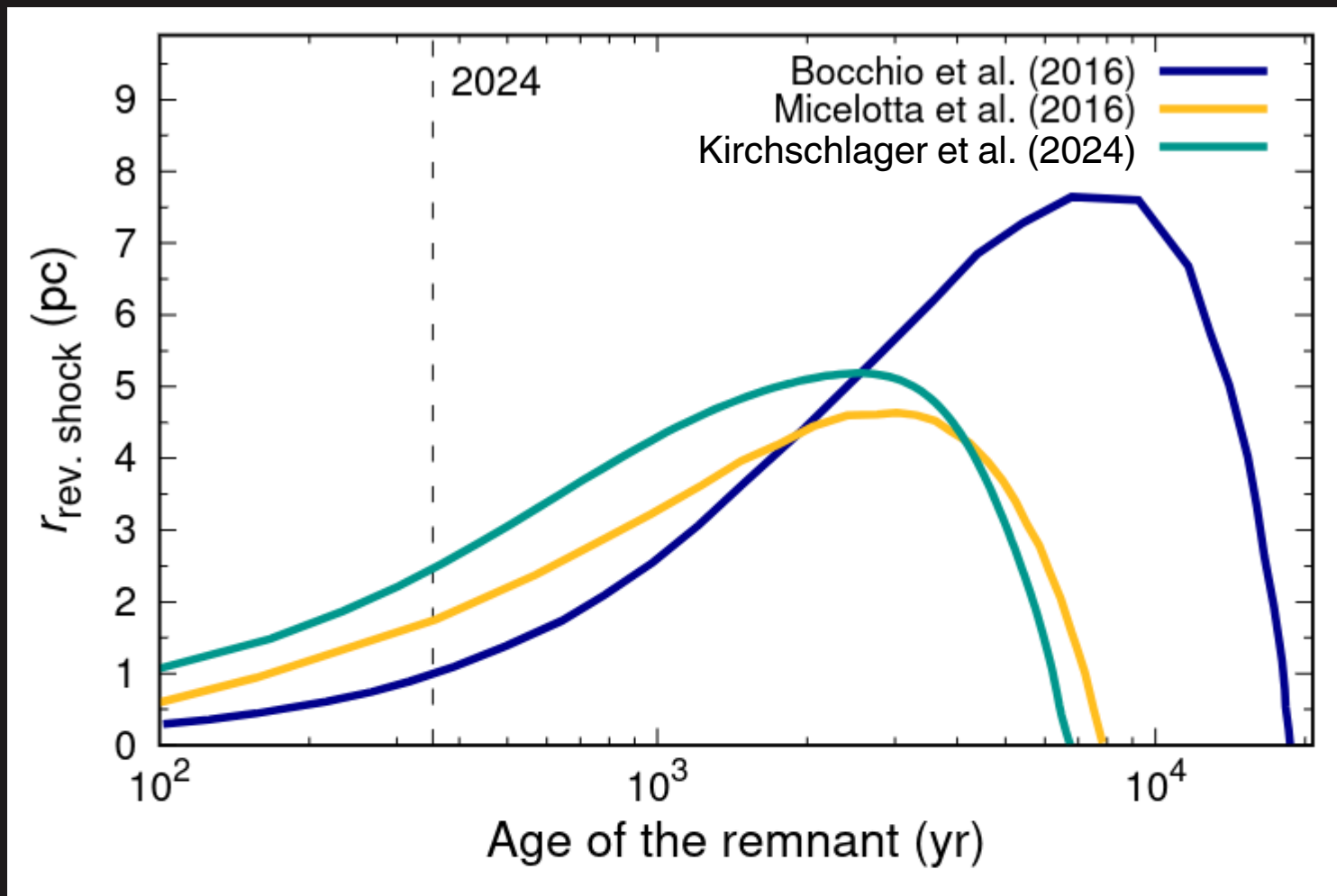


- Synthesis of the **XRISM-Resolve PV phase** spectrum using Orlando et al. (2020) MHD sim
- Spectra with largely **Doppler broadened emission lines**, due to the ejecta contribution



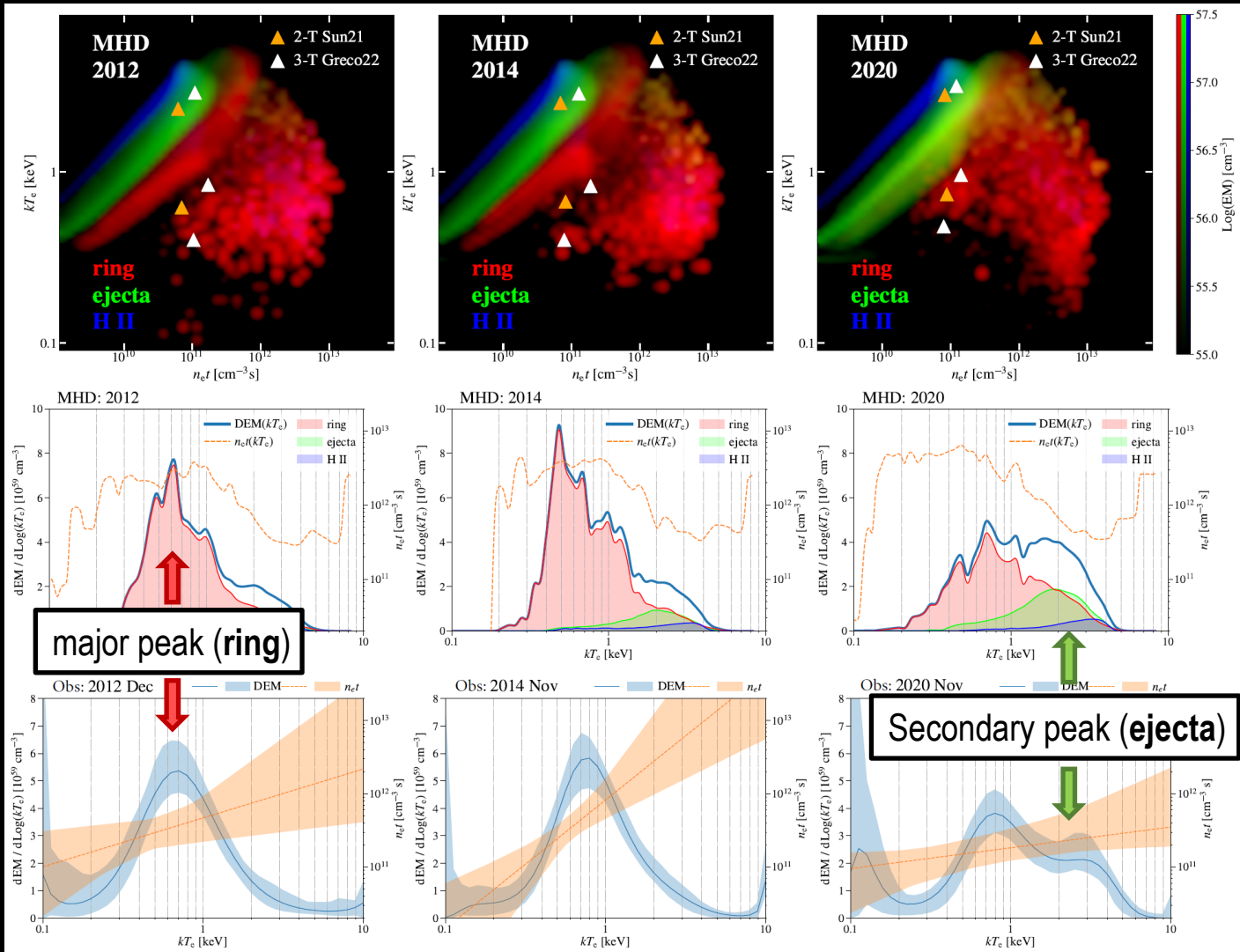
Truelove and Mckee:
RS there from the start

Our analytical solution:
RS formation time computed from ICs

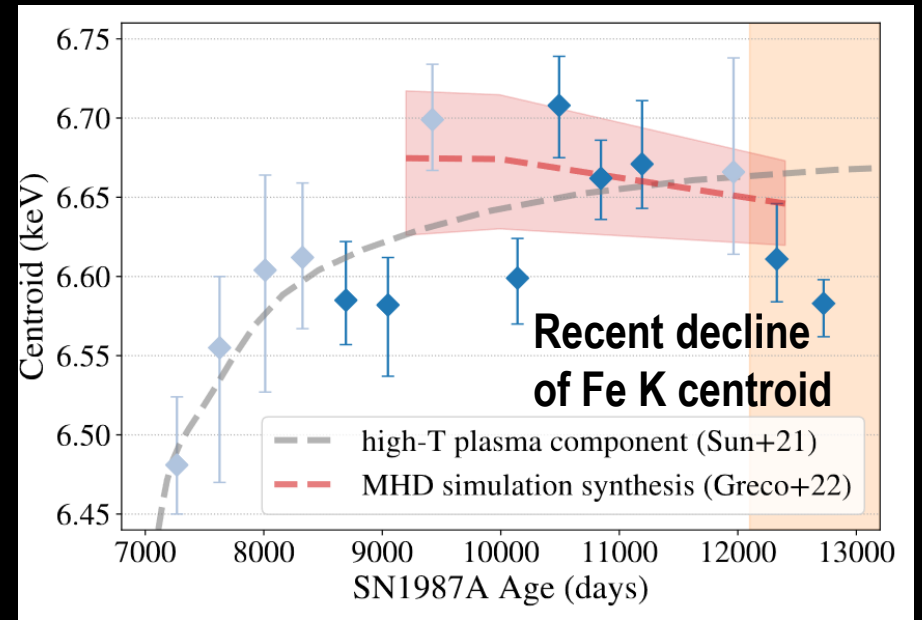


Differential Emission Measure Analysis of the X-ray Gas in SN 1987A: The Fading Ring & The Brightening Ejecta (S4.24)

Lei Sun, Emanuele Greco, Salvatore Orlando, Marco Miceli, Ping Zhou, Yang Chen, Jacco Vink



We performed a detailed DEM analysis of SN 1987A based on the long-term XMM-Newton observations. The recent emergence of a secondary peak in the DEM function together with the recent decline of the Fe K centroid energy indicate **a brightening ejecta component**.



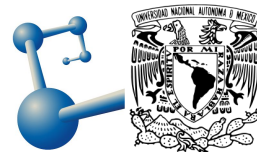
Simulated non-thermal emission of the supernova remnant G1.9+0.3

Villagran, et al. (2024). MNRAS 527, 1601-1611, DOI: 10.1093/mnras/stad322

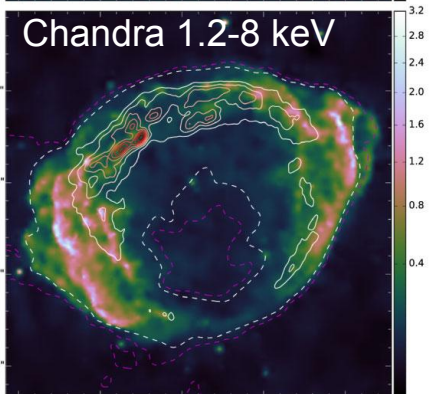
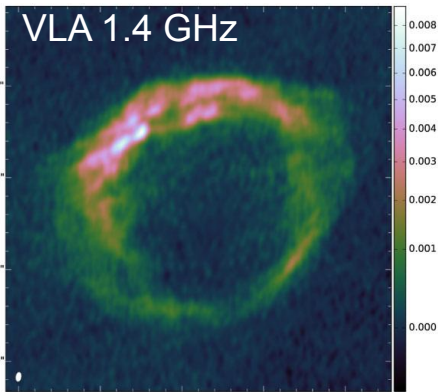
Presenter: J.C. Toledo-Roy, ICN-UNAM, Mexico City, juan.toledo@nucleares.unam.mx

SNRIII, 9-15 June

2024, Chania, Greece

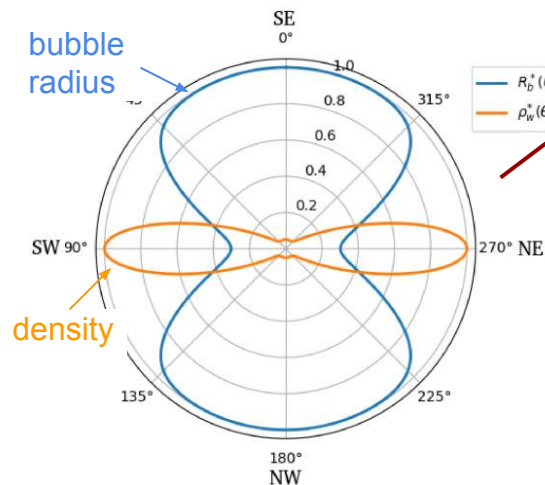


G1.9+0.3: a complex observed morphology



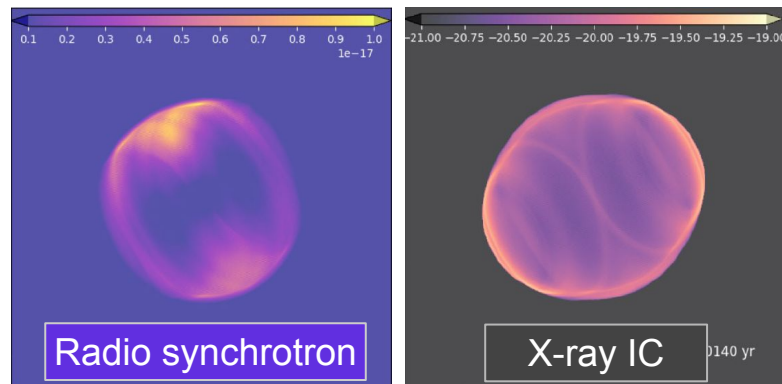
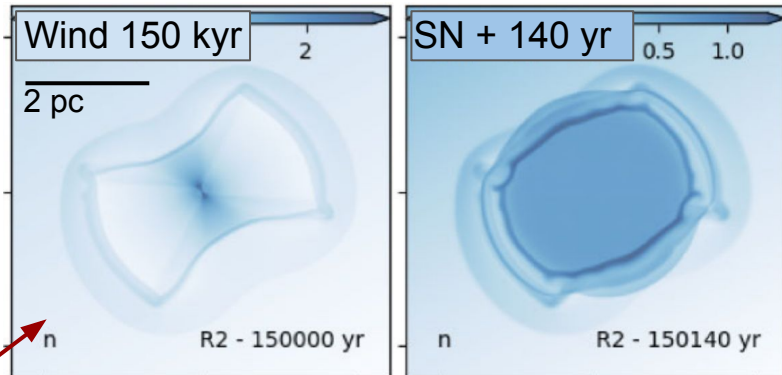
The model

- Pre-SN bipolar circumstellar nebula, either from companion or episodic PN phase
- Type Ia SN, isotropic
- Ambient galactic magnetic field
- Possible density gradient in ISM



Equatorially-dense anisotropic stellar wind, following Mellema et al. (1991)

MHD simulations and results



See the poster for further details!